

Outline

- What are land surface temperature (LST) and land surface (spectral) emissivity (LSE)?
- Hyperspectral LWIR airborne imaging instrumentation
 - NERC NCEO's OWL
 - NASA-JPL's HyTES
- LST + LSE retrieval from airborne hyperspectral imagers
 - HyTES's retrieval algorithm
- Validation of LST/LSE retrieval methods
 - Development of robust validation methodology
 - Assessment of HyTES retrieval algorithm
- Preliminary LST/LSE retrieval work OWL

Land Surface Temperature Mission



Copernicus High Priority Candidate (see Mike Perry's talk)

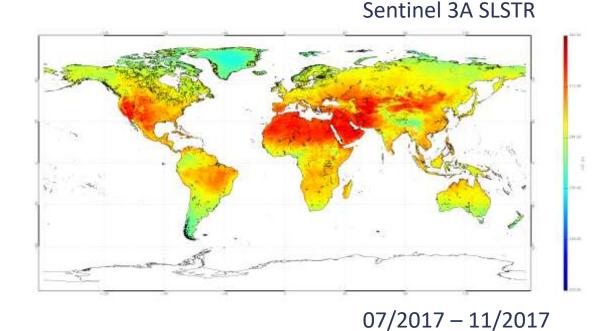




Land Surface Temperature (LST)

- Aggregated radiometric surface temperature of the ensemble of components within the sensor FOV
- Why important?
 - ✓ evaluating land surface & land-atmosphere interactions (e.g evapotranspiration)
 - ✓ constraining surface energy budgets (& model parameters)
 - ✓ providing observations of surface temperature change both globally and in key regions
- Estimated from TOA spectral radiance in Thermal Infrared atmospheric window (8 13 μ m)... but requires knowledge of other parameters

What sensor measures (spectral radiance)



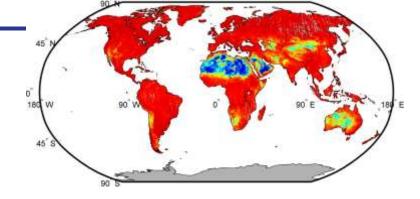
$$L_{\text{sen},\lambda} = \tau_{\lambda}(\theta) \left[\varepsilon_{\lambda} B_{\lambda}(\text{LST}) + (1 - \varepsilon_{\lambda}) L_{\text{sky},\lambda}^{\downarrow} \right] + L_{\text{sky},\lambda}^{\uparrow}(\theta)$$
emissivity

What we want to estimate

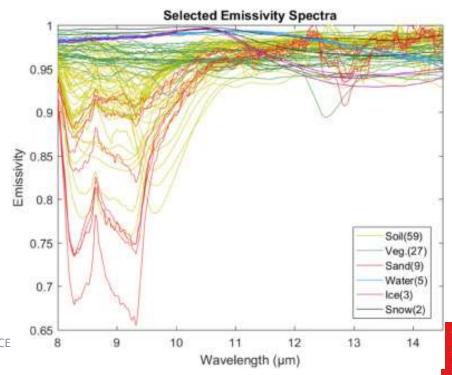


Land Surface [Spectral] Emissivity (LSE)

 Ratio of radiance emitted by objected to radiance that would be emitted by perfect emitter ('blackbody') at same temperature and wavelength



- Why important?
 - Calculating land surface temperature/ surface energy budgets
 - Land cover changes
 - Mineral mapping and resource exploitation
- Hyperspectral sensors offer new opportunity for simultaneous LST/emissivity retrieval + satellite mission development





Hyperspectral LWIR Airborne Instrumentation

- NCEO's Specim AisaOWL [OWL]
- NASA-JPL's Hyperspectral Thermal Emission Spectrometer [HyTES]







OWL	HyTES
Spectral range: 7.6 – 12.6 μm	Spectral range: 7.5 – 12.0 μm
96 spectral bands (50 nm bandwidth)	256 spectral bands (17.6 nm bandwidth)
TFOV = 24.2°	TFOV = 50.0°
At 1000m, pixel size 1.2m; swath ~410m (384 pixels)	At 1000m, pixel size 1.7m; (512 pixels)
Mass (scanhead): 13.1 kg	Mass (scanhead): 12 kg

Airborne data collected with HyTES in European sites (UK/Italy) June 2019 – data input for LSTM Design Studies



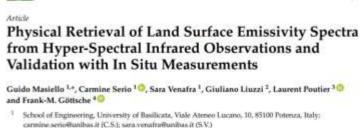


MDPI

LST/LSE retrieval

➤ Multiple algorithms been developed to tackle this





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remote sensing

Received: 17 May 2016; Accepted: 14 June 2018; Published: 20 June 2018



Abstract: A fully physical retrieval scheme for land surface emissivity spectra is presented,

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parameters, and

with high.

Temperature and Emissivity Retrievals From Hyperspectral Thermal Infrared Data Using Linear Spectral Emissivity Constraint

Ning Wang, Hua Wu, Françoise Nerry, Chuanrong Li, and Zhao-Liang Li

trace-Owing to the ill-posed problem of radiometric equathe separation of land surface temperature (LST) and land emissivity (LSE) from observed data has always been a esome problem. On the basis of the assumption that the pectrum can be described by a piecewise linear function, method has been proposed to retrieve LST and LSE from

Temperature-emissivity separation (TES), as one of the key problems in thermal infrared (TIR) remote sensing, is attractive, although many studies have been undertaken on this aspect [1]. However, the retrieval of LST and LSE from space is a difficult. task. First, the radiative transfer equation (RTE) shows that the pherically corrected by perspectral thermal infrared data us-





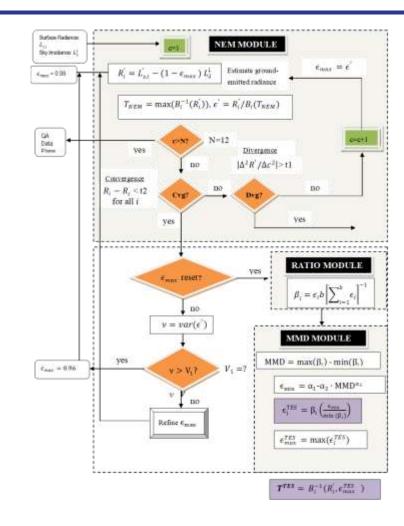
Transport Satellite (MTSAT), Given the different characteristics of the imagers onboard

Temperature and Emissivity Separation (TES) algorithm

> Combination of 3 different algorithms (NEM, Ratio, MMD)

Operational LST&E products using TES

- ASTER
- MODIS v6
- ASTER GED
- HyTES
- ECOSTRESS
- VIIRS (planned)









Validation of LST/ Emissivity Retrieval Algorithm







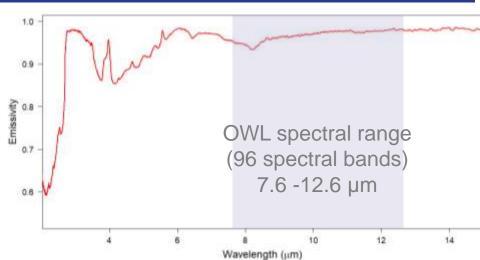


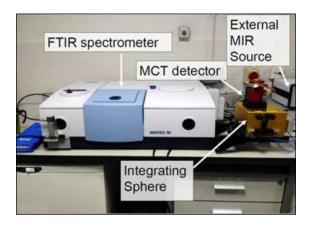
Temperature measurements from LWIR radiometers over thermally distinct surfaces

Spectral emissivity measurements from:

- (i) Samples collected + measured in laboratory
- (ii) Measurements in field using portable instruments











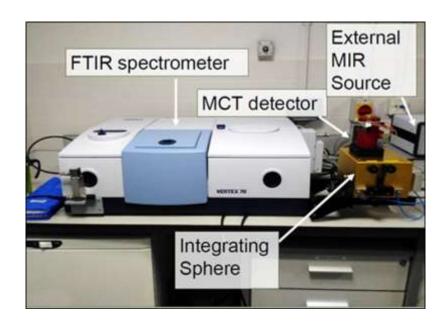
Laboratory Instrumentation: Emissivity

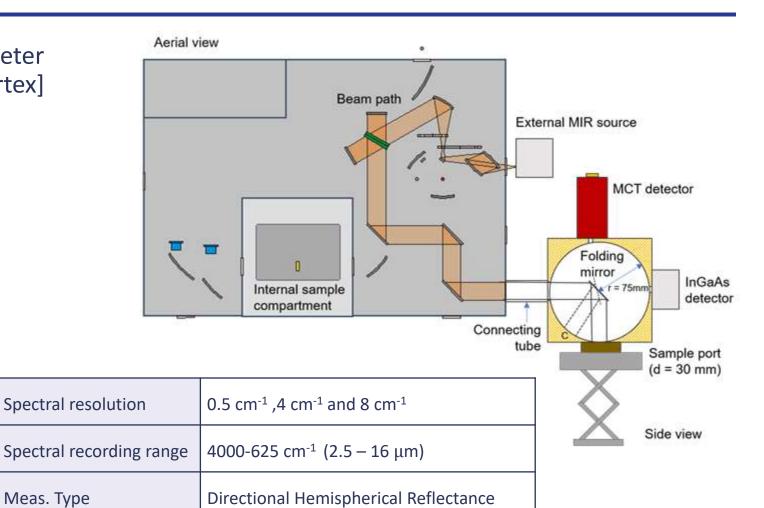
Meas. Type

Sample port

30 mm

➤ NCEO's Bruker Vertex V70 FTIR spectrometer with external gold integrating sphere [Vertex]









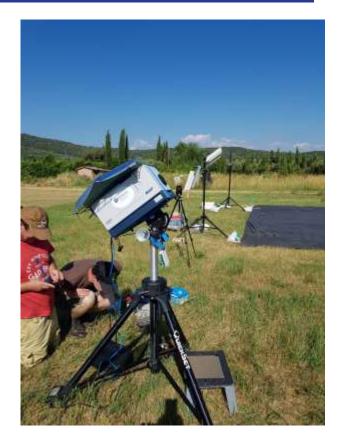
Field Instrumentation: Emissivity

- ➤ NCEO's Bruker EM27 Open Path FTIR spectrometer [EM27]
- ➤ NASA-JPL's Designs & Prototypes microFTIR spectrometer [D&P]

	EM27	D&P
Spectral resolution	0.5 cm ⁻¹ ,4 cm ⁻¹	6 cm ⁻¹
Spectral recording range	5000 – 700 cm ⁻¹ (2 – 14 μm)	3333 - 2000 cm ⁻¹ (3 - 5 μm); 1250 – 833 cm ⁻¹ (8 - 12 μm)
Туре	Passive Emission	Passive Emission
FOV at 1m	60 mm	80 – 160 mm (depending on foreoptics)
Mass/ Power	18 kg, 40 – 80 W	12.5 kg, 18 W



EM27 measuring LWIR surface emissivity



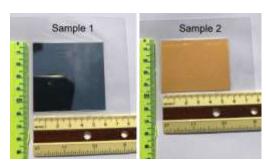
EM27 measuring LWIR downwelling irradiance



Can we trust these 'truths'? - Laboratory Round Robin

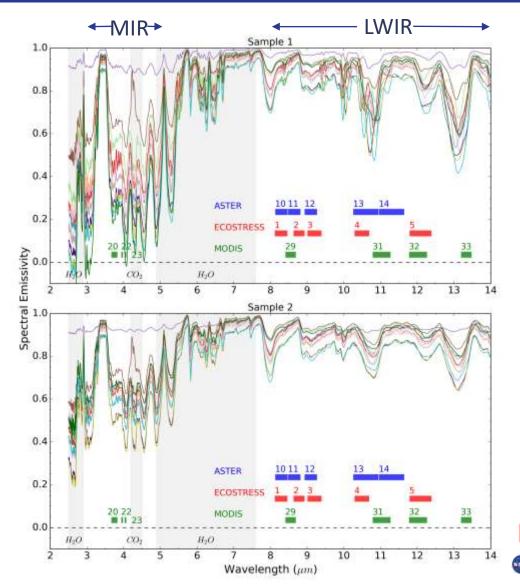
Intercomparison of measurements from **13** different setups at **8** laboratories (incl NCEO-King's, NASA JPL, DLR..)

Samples: aluminium/gold sheets laminated in polyethylene



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Standard Deviation over LWIR (% mean)

Sample 1a: 0.142 (16.6%) Sample 2a: 0.110 (12.5%)

Higher uncertainties from laboratory measurements of emissivity than previously assumed





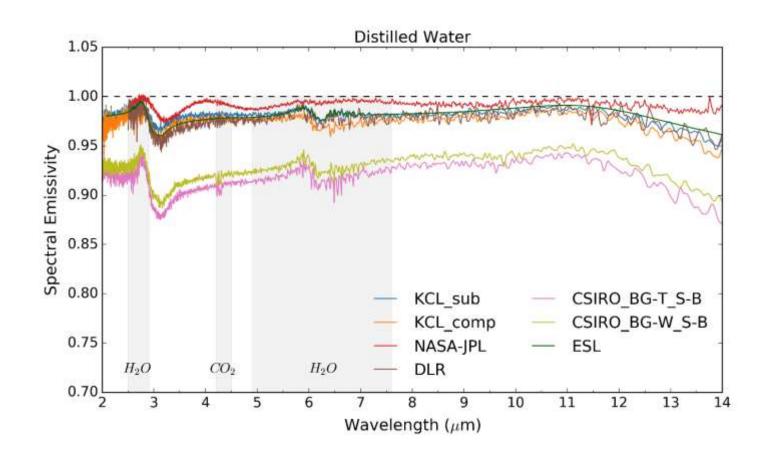








Can we trust these 'truths'? - Laboratory (2)



Differences observed amongst measurements of distilled water

Use of lowest measurement would result in LST 2.9 K less than if used highest emissivity

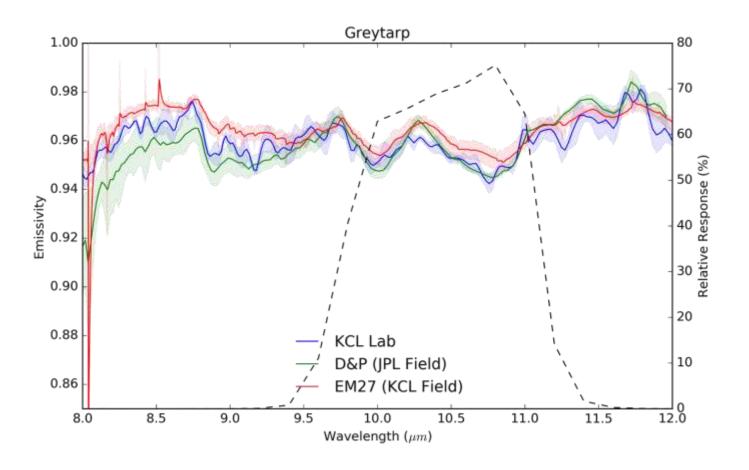
Amongst higher emissivity group, differences would lead to surface temperature retrieval differences of 0.7 K







Can we trust these 'truths'? - Field/Laboratory









Method	Heitronics KT15.85 band-specific emissivity
Laboratory (Vertex-V70)	0.956 ± 0.003
Field (EM27)	0.952 ± 0.009
Field (D&P)	0.956 ± 0.002





Data Collection [HyTES] during ESA/NASA NETSense Campaign 2019

Data collected as part of NETSense campaign (June 2019)





Grosseto, Italy







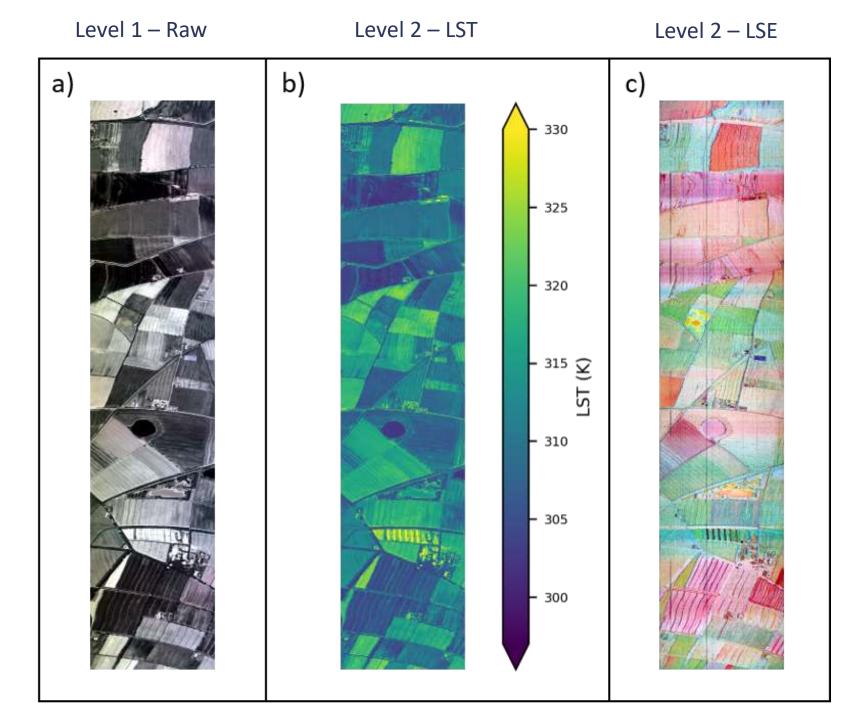


Duxford + surrounding areas, UK

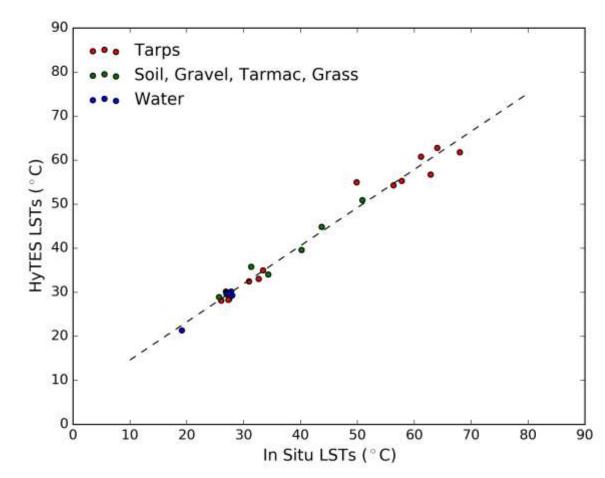


HyTES data from NETSense Campaign

Grosseto AM 23 June 2019



HyTES Airborne vs. In Situ LST Data Comparison

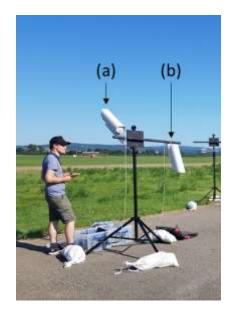


All surfaces:

Bias = + 1.35 °C Scatter = 2.21 °C

Just land:

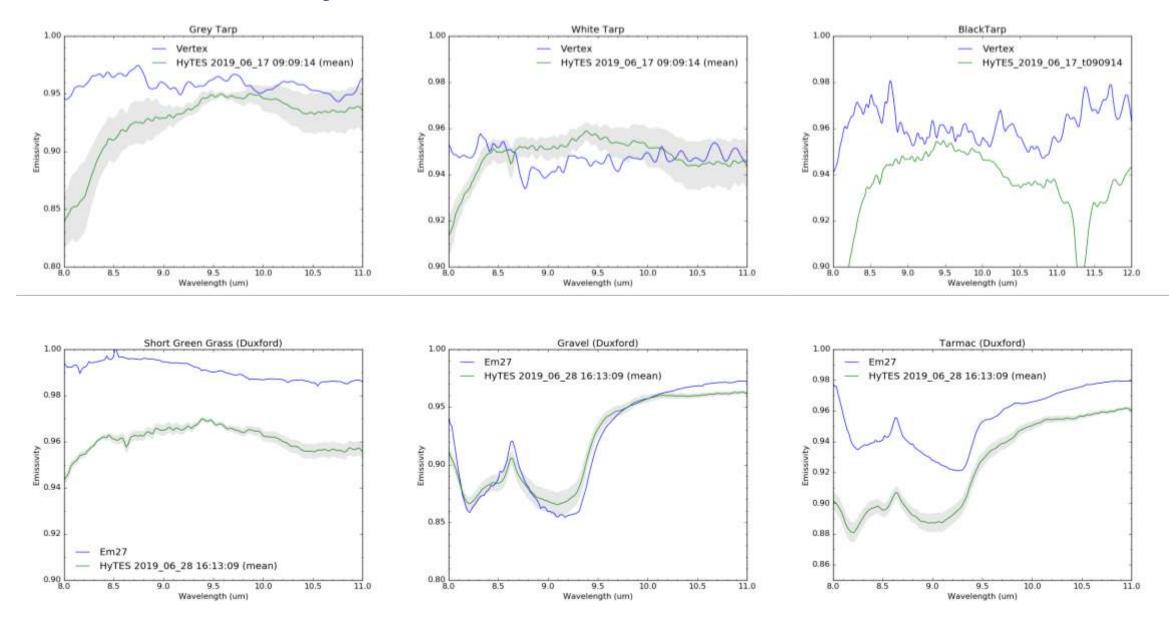
Bias = + 0.8 °C Scatter = 2.21 °C







HyTES – In Situ Field/Lab LSE



Data Collection (OWL)



R: 8.03μm G: 10.0μm B: 12.0μm





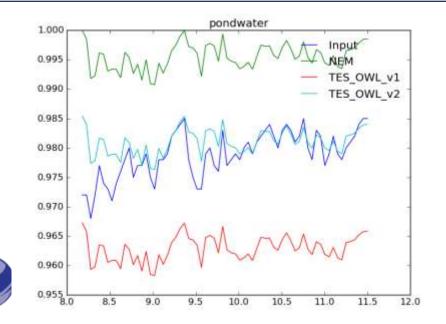




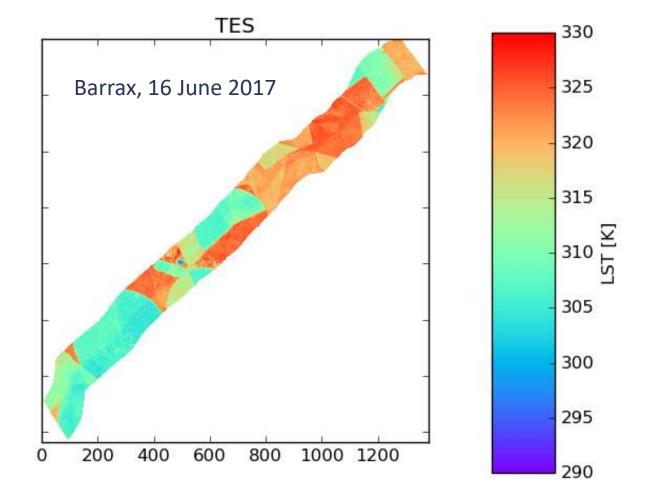
LST/Emissivity algorithm development: OWL Airborne Sensor

(1) Testing HyTES alg. adapted to OWL on simulated data

Sample	LST bias [OWL_v1] (K)	LST bias [OWL_v2] (K)
Water	1.096	0.161
Soil	1.966	1.125
Rock	0.338	0.734



(2) Testing OWL-derived LSTs and emissivity vs. in situ data



Summary and Concluding Remarks

- Hyperspectral airborne sensors offer new opportunities for mission and alg development for LST/LSE
- HyTES' LST and LSE retrieval alg. has been tested through deployment of field and lab instrumentation
 HyTES LSTs found to be within 1.35 K for all surfaces considered and 0.8 K for natural surfaces
- Evaluation of algorithms must take into account accuracy of field/laboratory instrumentation and outputs o Intercomparison of different laboratory emissivity setups suggests NCEO laboratory setup within 2% of mean over LWIR
- Early application of HyTES LST/emissivity retrieval algorithm to OWL data promising
 OWL algorithm within 1.2 K when tested on simulated data of natural surfaces
- Next steps: OWL validation with existing in situ data and HyTES 2021 campaign in Barrax flying OWL alongside in same platform



